

# Snowmass2021 - Letter of Interest

## *Studies of the Muon Excess in Cosmic Ray Air Showers*

### Thematic Areas:

- (CF7) Cosmic Probes of Fundamental Physics
- (EF6) QCD and Strong Interactions: Hadronic Structure and Forward QCD
- (EF7) QCD and Strong interactions: Heavy Ions
- (AF4) Multi-TeV Colliders

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**Abstract:** Observations of cosmic rays with PeV energies and higher rely on indirect measurements of extensive air showers (EAS) in the Earth’s atmosphere. The inference of the properties of the initial cosmic ray requires precise physical models of the air shower. Evidence for an excess of GeV muons in EAS with respect to model predictions has been reported by many experiments over the last 20 years. This excess was recently confirmed in a meta analysis of measurements from eight air shower experiments, for all currently available hadronic interaction models. This letter of interest focuses on the current status of muon measurements from observations of EAS over four orders of magnitude in cosmic ray energy and their discrepancies with respect to predictions from recent hadronic interaction models. We will address the main challenges in the description of muon production in EAS and identify measurements that are required to improve current model predictions. We will further discuss ongoing studies and future perspectives for muon measurements, and emphasize the importance of synergies between air shower observations and dedicated collider measurements of nucleus-nucleus interactions in the forward region in order to understand the origin of the muon excess in EAS.

## Introduction

Cosmic rays have been measured in the Earth's atmosphere with energies exceeding  $10^{20}$  eV, but their sources remain unclear, their acceleration mechanisms and mass composition are uncertain, and several features observed in the energy spectrum are not well understood<sup>1</sup>. Large ground-based experiments are used to observe high energy cosmic rays indirectly through extensive air showers (EAS) that they generate in the atmosphere. The inference of cosmic ray properties requires a precise physical model of the air shower, but key properties of the EAS development are driven by forward hadron production in soft and semihard hadron-ion collisions, which cannot be computed *ab initio*, and their modeling is subject to large theoretical uncertainties. In fact, none of the current hadronic interaction models tuned to available collider data is able to correctly describe muon production in EAS<sup>2,3</sup>, a problem commonly referred to as the *muon puzzle* in air shower physics<sup>4</sup>.

## The Muon Excess in Extensive Air Showers

An excess of muons with respect to simulated air showers was first experimentally observed in the year 2000<sup>5</sup>. Subsequently, the excess was qualitatively confirmed by several experiments<sup>6–9</sup>, while other experiments reported no excess<sup>10–13</sup>. The most unambiguous experimental evidence of the excess was revealed in the analysis of Auger data<sup>14,15</sup>. As hadronic interaction models were further developed and tuned to more data, in particular from LHC and SPS, their predictions for air showers improved but discrepancies in predicted muon production remained. Recently, a global picture of the energy dependence and high statistical significance of the muon excess was found using measurements of GeV muons from eight air shower experiments<sup>8–16</sup>, ranging over more than four orders of magnitude in shower energy, by a multi-collaborative *Working Group on Hadronic Interactions and Shower Physics* (WHISP)<sup>2,3</sup>. Key aspects of this analysis are a cross-calibration of the energy scales of the different experiments and the definition of the  $z$ -value, an abstract measure of the muon content which is comparable between experiments and different analyses. The value  $\Delta z = z - z_{\text{mass}}$ , where  $z_{\text{mass}}$  is the number of muons predicted by a hadronic model assuming a mass composition of the primaries based on experimental parameterizations<sup>17</sup>, measures the difference between the experimental data and the inferred number of muons for a given hadronic interaction model. A positive value indicates an excess of muons in data with respect to simulations and 0 indicates a perfect match. The resulting distribution of  $\Delta z$  is shown in Fig. 1 for two hadronic models, EPOS-LHC<sup>18</sup> and QGSJet-II-04<sup>19</sup>. The experimental measurements are consistent within uncertainties with predictions up to energies of approximately 100 PeV, corresponding to proton-air collisions at nucleon-nucleon center-of-mass energies of  $\sqrt{s_{NN}} \lesssim 14$  TeV. However, above 100 PeV a muon excess is observed that systematically increases with the shower energy. The slope of a linear fit to this excess is found to be non-zero at  $\sim 8\sigma$  for all hadronic interaction models considered.

Discrepancies in measurements of the attenuation length of GeV muons<sup>12</sup> suggest a problem in the description of the energy spectra of muons at production<sup>20–22</sup>. Preliminary studies also indicate differences in

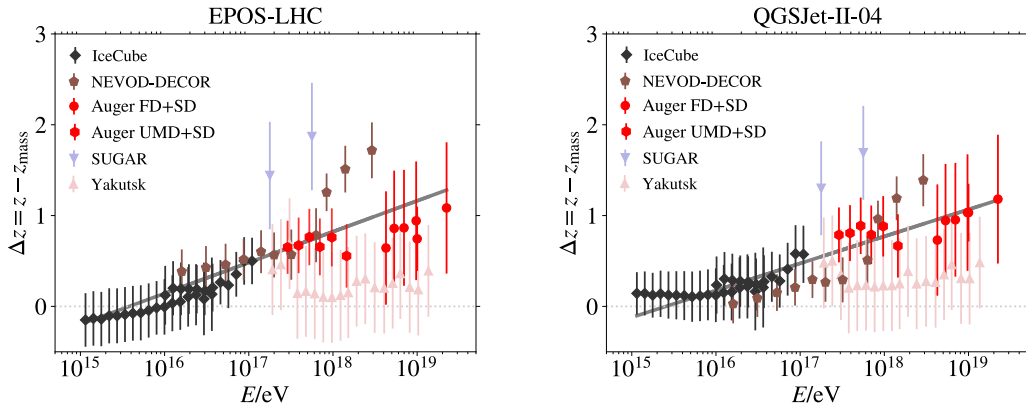


Figure 1: Distribution of  $\Delta z = z - z_{\text{mass}}$  from various experiments and corresponding linear fit for two hadronic interaction models, EPOS-LHC and QGSJet-II-04 (see WHISP report<sup>2</sup> for details).

the discrepancies of GeV and TeV muons<sup>23</sup> and inconsistencies in the zenith angle distributions of simulated TeV muons have been reported<sup>24–26</sup> which are not yet understood. The muon production depth in the atmosphere<sup>27,28</sup> has also been shown to have strong discrepancies with respect to models, which has been linked to a poor description of pion-air interactions in EAS.

## Upgrades to EAS Detectors and Synergies with Accelerator Measurements

The main challenge in the description of muon production in the atmosphere is the treatment of hadron-ion interactions in the forward region over many orders of magnitude in energy<sup>22,29</sup>. Conventional options to improve the description of muon production in hadronic models are the enhancement of the production of baryon-pairs, neutral rho mesons<sup>30</sup>, or heavy quarks<sup>31</sup>, as well as possibly missing physics in the soft-QCD regime of hadron-ion collisions<sup>31–34</sup>, or other new phenomena<sup>35–37</sup>.

Various analyses of atmospheric muons in the GeV range are in preparation<sup>38–41</sup>, using improved analysis methods and the most recent hadronic models. Analyses based on global fits of different shower observables have been developed which will provide reduced uncertainties<sup>42–45</sup> and machine learning techniques will further enhance the capabilities of existing detectors<sup>46,47</sup>. Ongoing detector upgrades of existing facilities, such as AugerPrime<sup>48–50</sup> and IceCube-Gen2<sup>51–55</sup>, will enhance the precision of air shower measurements and reduce uncertainties in the interpretation of muon data which arise from energy scale or mass composition uncertainties. Upgrades of the NEVOD-DECOR experiment<sup>56,57</sup> will improve the measurement of the energy deposition of muons which will also help to test and constrain hadronic model predictions. It has also been shown that the shower-to-shower fluctuations of the number of muons are strongly dominated by the first interaction of the EAS<sup>58,59</sup>. Thus, they are a direct probe of the hadron production spectra and provide a test of multiparticle production in hadron-ion interactions at center-of-mass energies above those reachable at the LHC. IceCube offers the unique opportunity to measure the EAS muon content simultaneously at two vastly different energy scales, in the deep ice detector (TeV muons), as well as with IceTop<sup>23,60</sup> and new detectors<sup>52,53</sup> at the surface (GeV muons), which will test and strongly constrain muon enhancement models<sup>30</sup>.

Ongoing studies on the muon discrepancy at the LHC include detailed measurements of prompt hadron production cross-sections in the forward region in proton-proton and proton-lead collisions and measurements of the ratio of electromagnetic and hadronic energy flows, to which the production of GeV muons is particularly sensitive<sup>34,61–63</sup>. Similar studies at central rapidities improved the hadronic models with the first LHC data<sup>64</sup>. Running the LHC with oxygen beams is planned for 2023<sup>62</sup>, which was largely motivated by the muon excess and the need to study the nuclear modification of hadron production in the first collisions in EAS. The LHCb SMOG-2 upgrade<sup>65</sup> will extend the unique capability of LHCb to perform LHC fixed target experiments to study late interactions in EAS, and ALICE will extend LHC fixed-target measurements to a wider rapidity range<sup>66</sup>. Data from the FASER<sup>67</sup> and FASER $\nu$ <sup>68,69</sup> experiments, measuring forward muon and neutrino fluxes at the LHC, as well as from AFTER@LHC<sup>70</sup> and LHCf<sup>71</sup>, will also provide useful information. Further details of the synergy between air shower and collider physics are discussed in a separate letter of interest<sup>72</sup> for Snowmass2021.

## Conclusions

A meta analysis of muon data from multiple experiments recently confirmed an excess of atmospheric muons in EAS. Some observations also suggest a problem in the modeling of muon production spectra but a possible connection to the physics of the excess remains unknown. Updates of analyses from existing and new air shower facilities, which include measurements of the muon production depth, the muon energy deposit, and measurements of muons at multiple energies, as well as shower-to-shower fluctuations, will provide crucial information on hadron-ion interactions beyond the phase space of existing colliders in order to constrain muon production models. Dedicated collider measurements will provide complementary information of prompt hadron production in the forward region. This synergy between air shower and collider experiments will be a crucial step towards an understanding of the origin of the muon discrepancies in EAS.

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